

# ● Berömda Experiment

- |      |                                  |             |
|------|----------------------------------|-------------|
| 1933 | Uppträckten av positronen        | → boken 1.2 |
| 1968 | Uppträckten av kvarkar           | → boken 7.4 |
| 1973 | Uppträckten av neutral ström     | → boken 8.0 |
| 1974 | Uppträckten av J/ψ               | → boken 3.4 |
| 1979 | Uppträckten av gluonen           | → boken 7.2 |
| 1983 | Uppträckten av $Z^0$ och $W^\pm$ | → boken 8.1 |

## Uppräckten av positronen (1933)

Laboratorium = California Institute of Technology

Accelerator = Universum ( $E_{beam} = 1 - 10^8 \text{ GeV}$ )

Experiment = Kosmiskt dimkammerexperiment

Processer = Tex  $p + O \rightarrow \pi\gamma^+ + X$   
                  ↳  $\mu^+ + \nu_\mu$   
                  ↳  $e^+ + \nu_e + \bar{\nu}_\mu$

Nobelpristagare = Anderson

# Kosmisk stråling

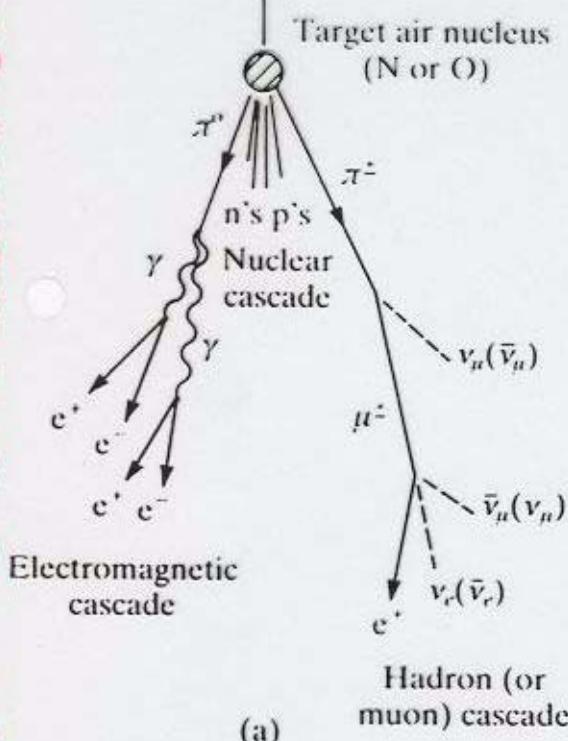
p. Fe, etc. (or  $\gamma$ )

Supernovor, Pulsser, Kvarnar...

$\gamma, e^\pm, p, \bar{p}, Fe, v...$

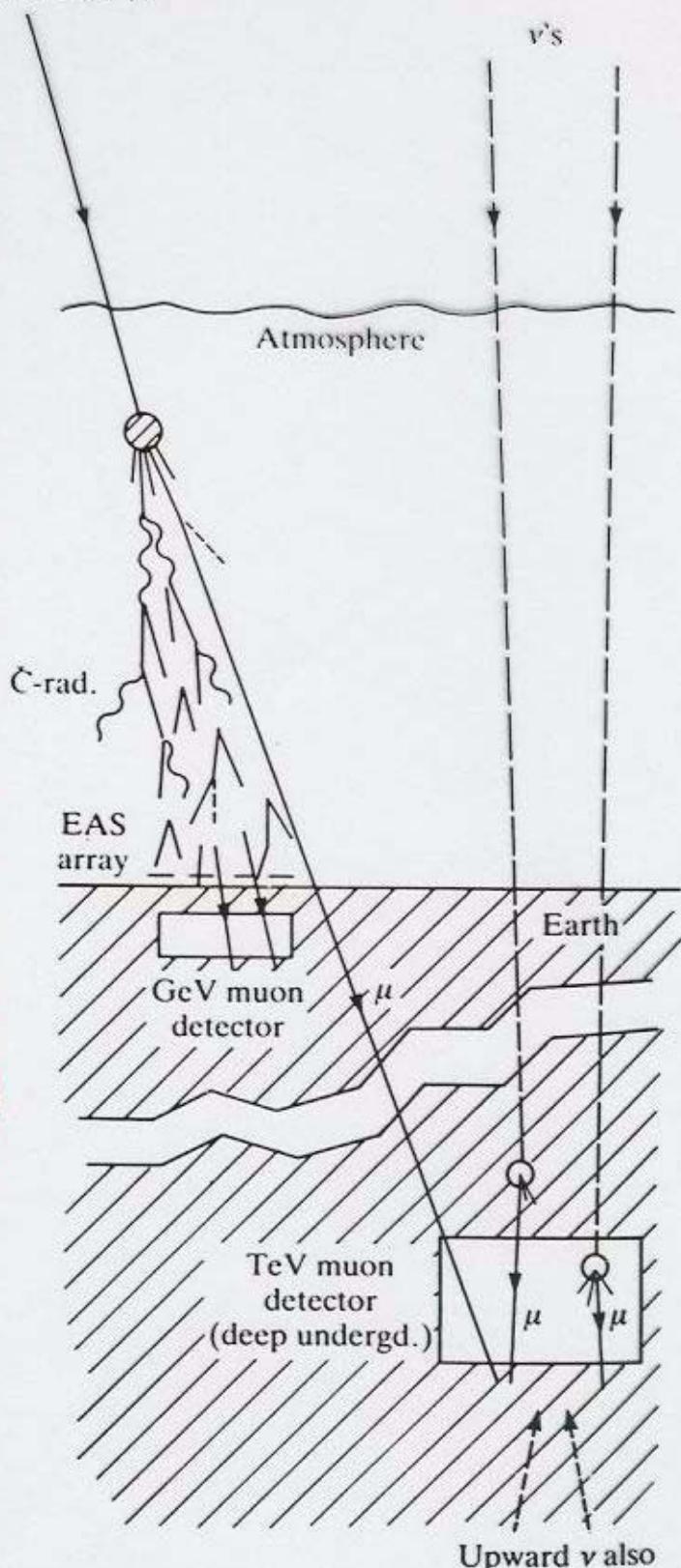
Primary p or Fe etc.

$E = 1 - 10^{10} \text{ GeV}$



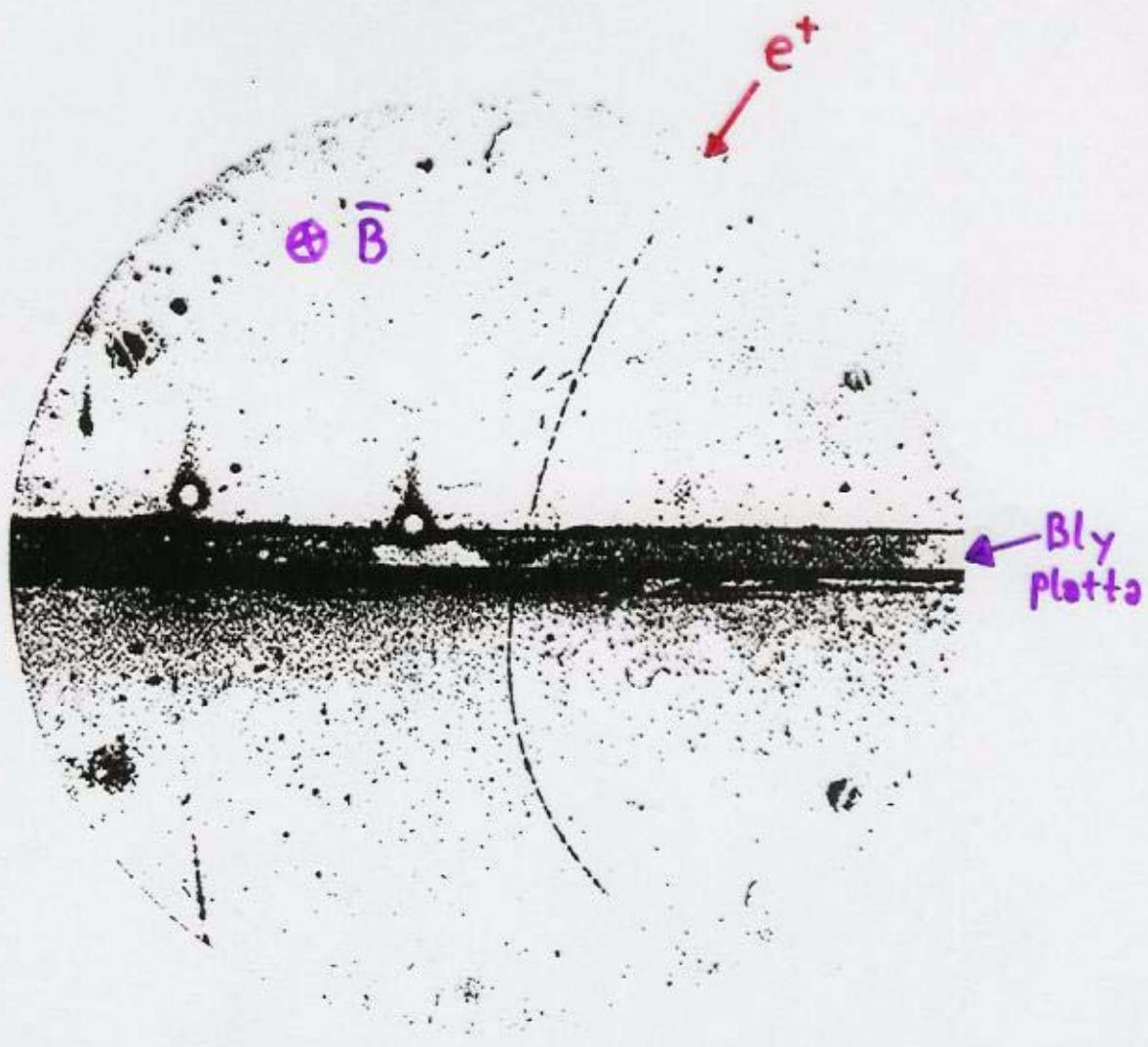
EAS : Extensive Air Shower

(b)



**Fig. 18.3** (a) The development of an air shower caused by a primary proton or heavy nucleus hitting an air nucleus, and producing a nuclear cascade of neutrons and protons, as well as pions which subsequently decay and produce electromagnetic cascades (via  $\pi^0 \rightarrow \gamma\gamma$ ) or so-called "hadron" cascades through  $\pi^\pm \rightarrow \mu^\pm \nu$ . (b) A summary of the techniques used to observe cosmic ray particles. The Čerenkov radiation that accompanies the cascade can be detected on Earth by Čerenkov telescopes even when the cosmic-ray shower does not penetrate to the ground.

## Dimkammerbild zu ein positron



# Uppräckten av "neutral current" (1973)

Laboratorium = CERN

Accelerator = PS

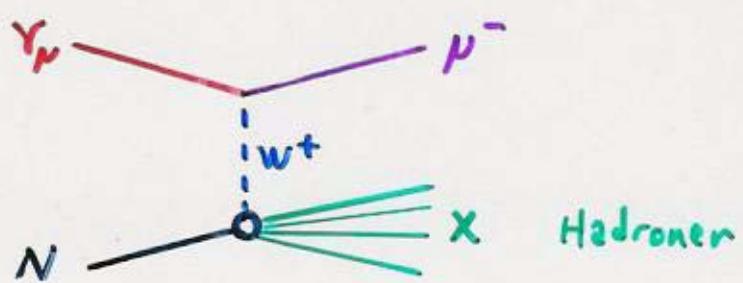
Experiment = Gargamelle ( $\nu N$  fixed target)

Processor =  $\begin{array}{l} \nu_\mu + N \rightarrow \nu_\mu + X \\ \bar{\nu}_\mu + N \rightarrow \bar{\nu}_\mu + X \end{array} \left. \right\}$  Neutral current

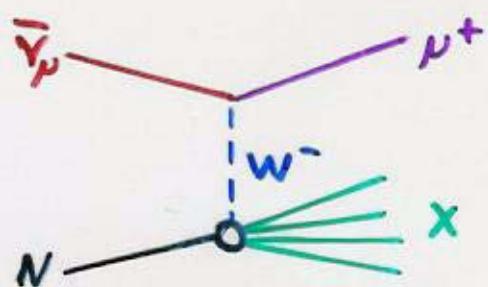
$\begin{array}{l} \nu_\mu + N \rightarrow \mu^- + X \\ \bar{\nu}_\mu + N \rightarrow \mu^+ + X \end{array} \left. \right\}$  Charged current

## Charged current:

$$\nu_\mu + N \rightarrow \mu^- + X$$

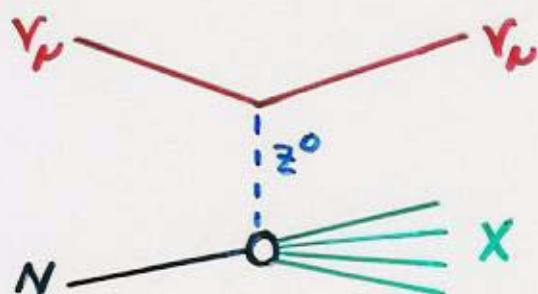


$$\bar{\nu}_\mu + N \rightarrow \mu^+ + X$$

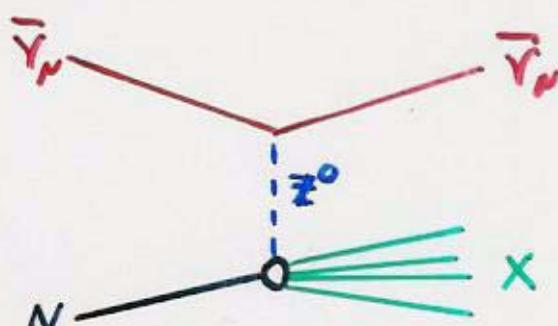


## Neutral current:

$$\nu_\mu + N \rightarrow \nu_\mu + X$$

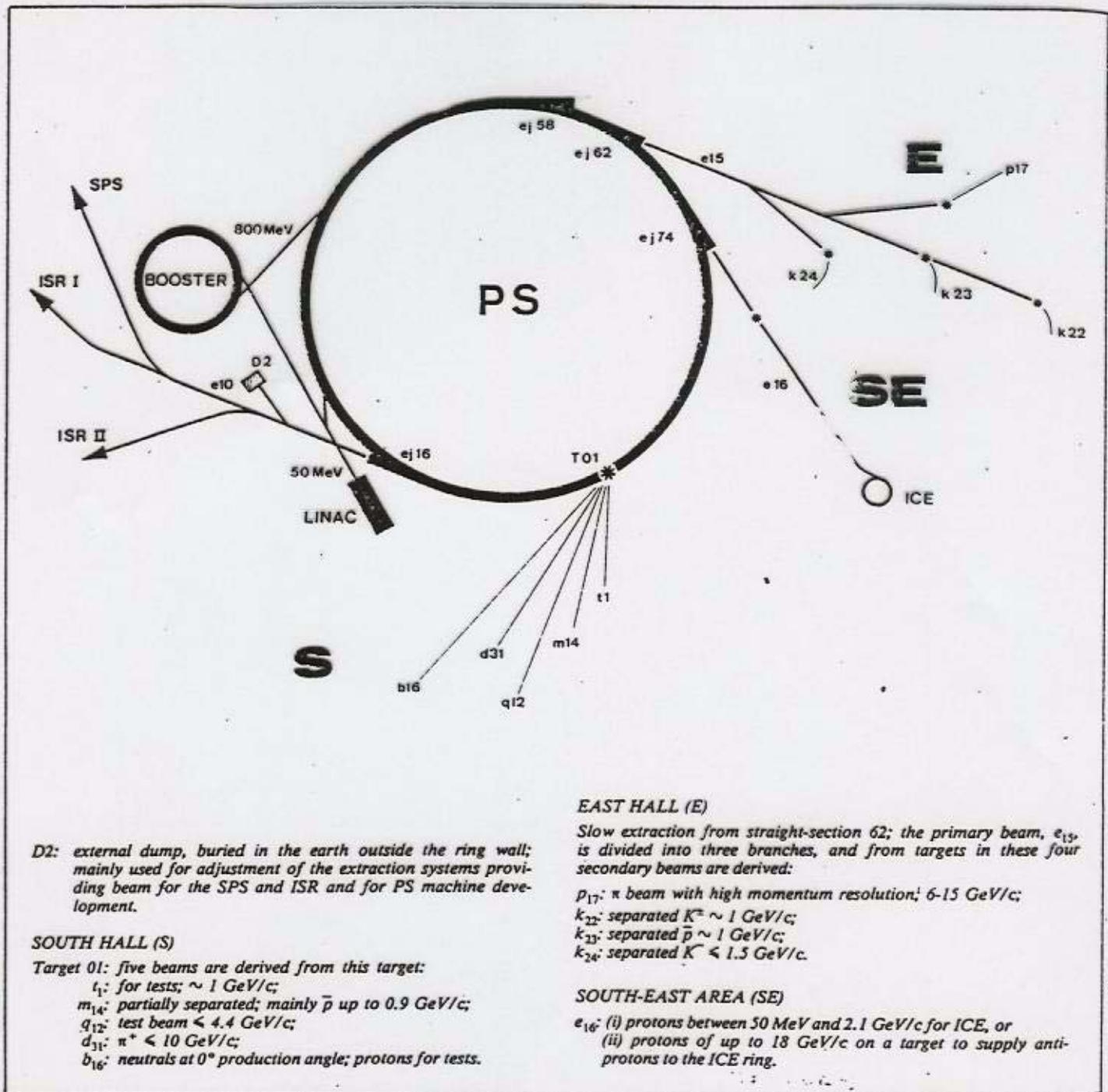


$$\bar{\nu}_\mu + N \rightarrow \bar{\nu}_\mu + X$$

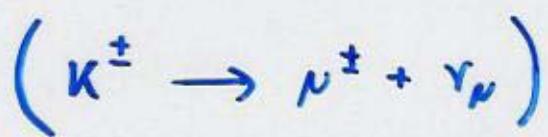
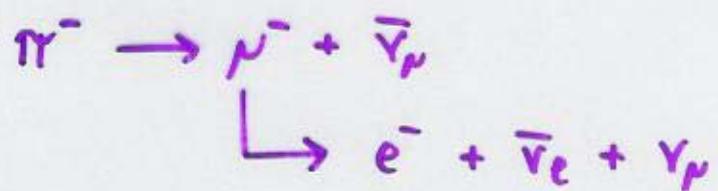
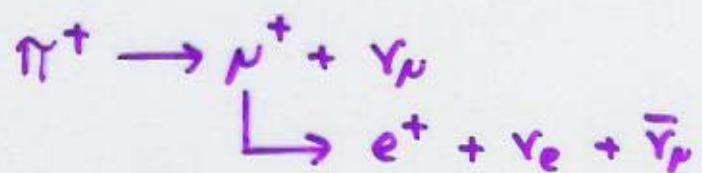
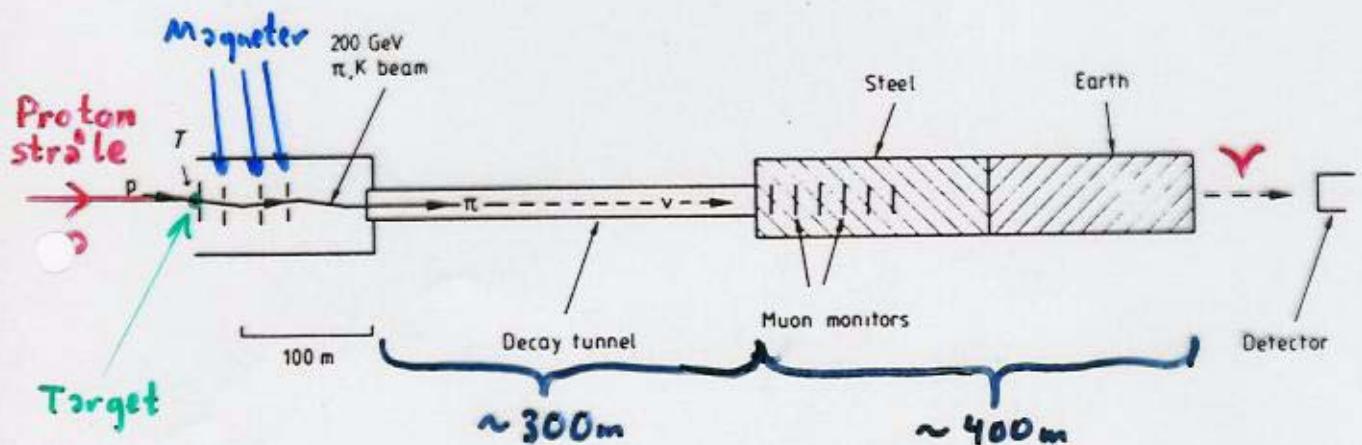


OBS: En neutrino kan ej växelverka elektromagnetiskt  
dvs genom fotonutbyte.

# PS - Accelerator

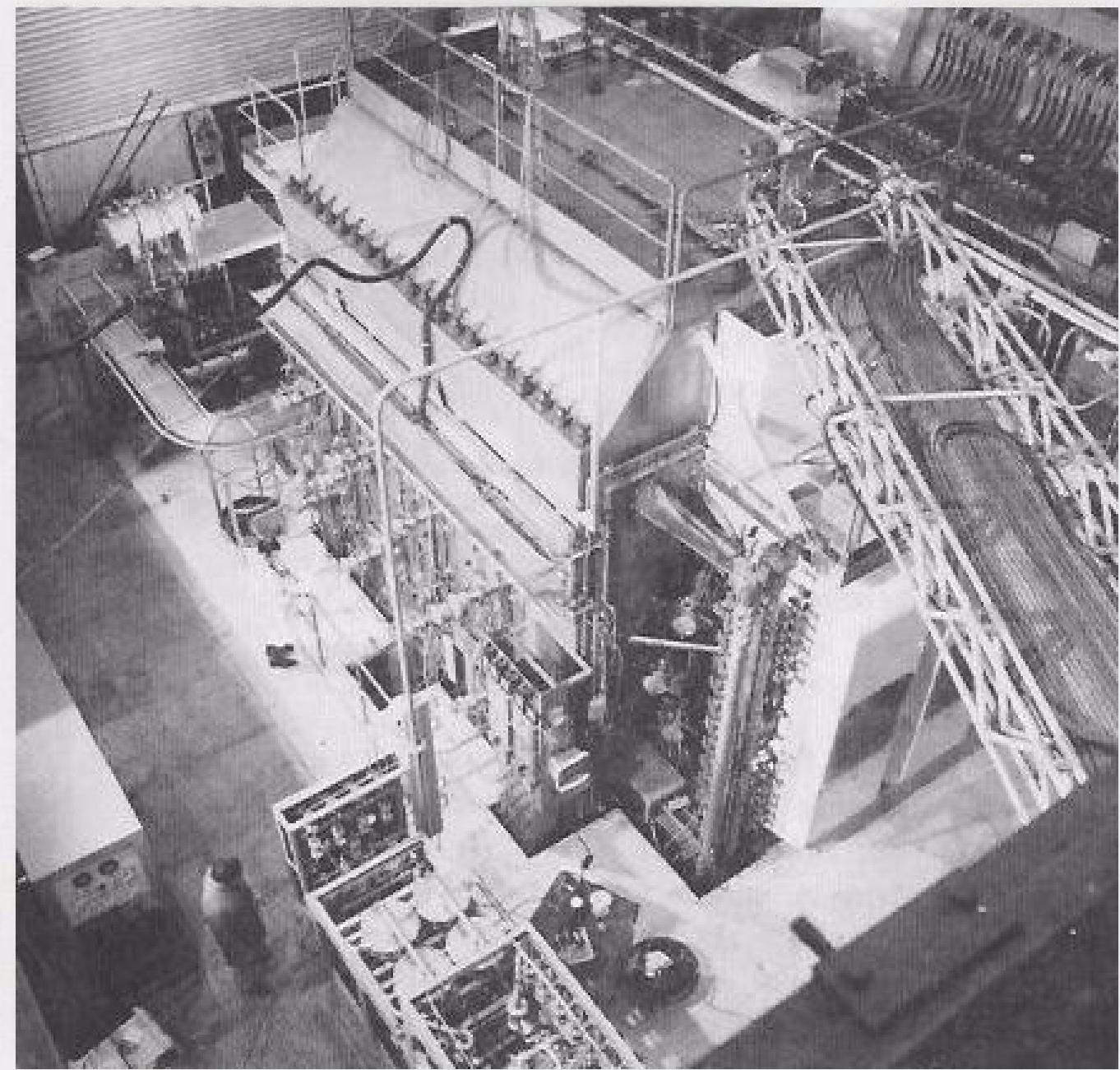


## Tillverkning av en neutrinostråle



# GARGAMELLE

**6.4** Bubbelkammaren GARGAMELLE vid CERN (utsidan). Kammaren är fylld med en speciell vätska. Så fort en stråle med neutriner far igenom kammaren reducerar man plötsligt trycket i kammaren. De laddade partiklarna bildar spår i kammaren bestående av blåsor, vilka fotograferas. Kammaren är omgiven av en magnet, som skapar ett starkt magnetfält i kammarens inre. Av riktning och styrka på de laddade partiklarnas avböjning i magnetfältet kan man dra slutsatser om deras energi och laddning.



## Bubbelkammarbild av en "neutral ström" händelse

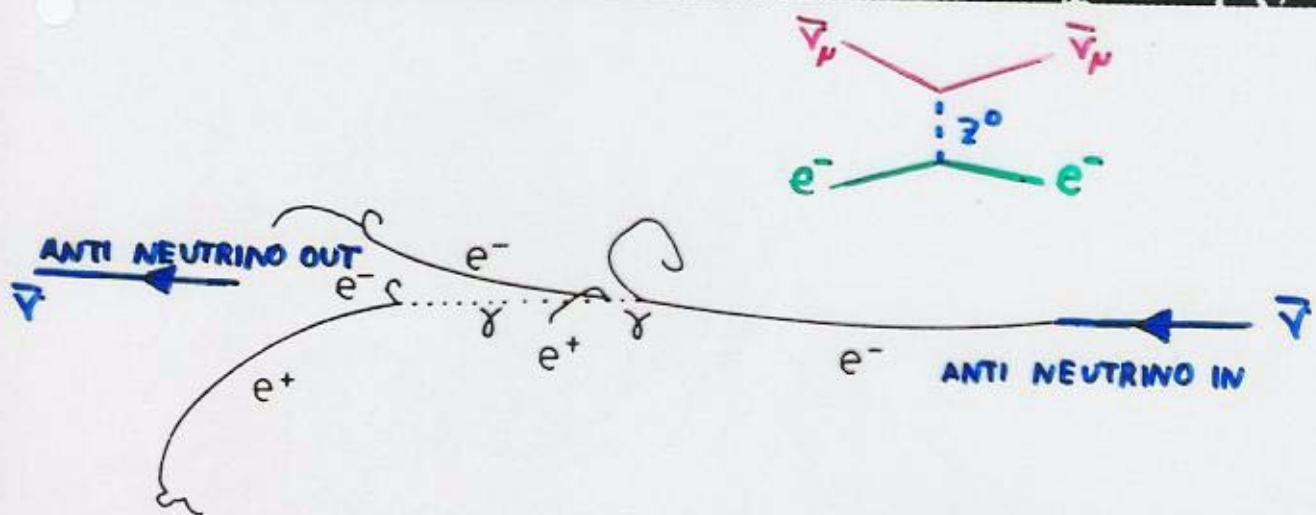
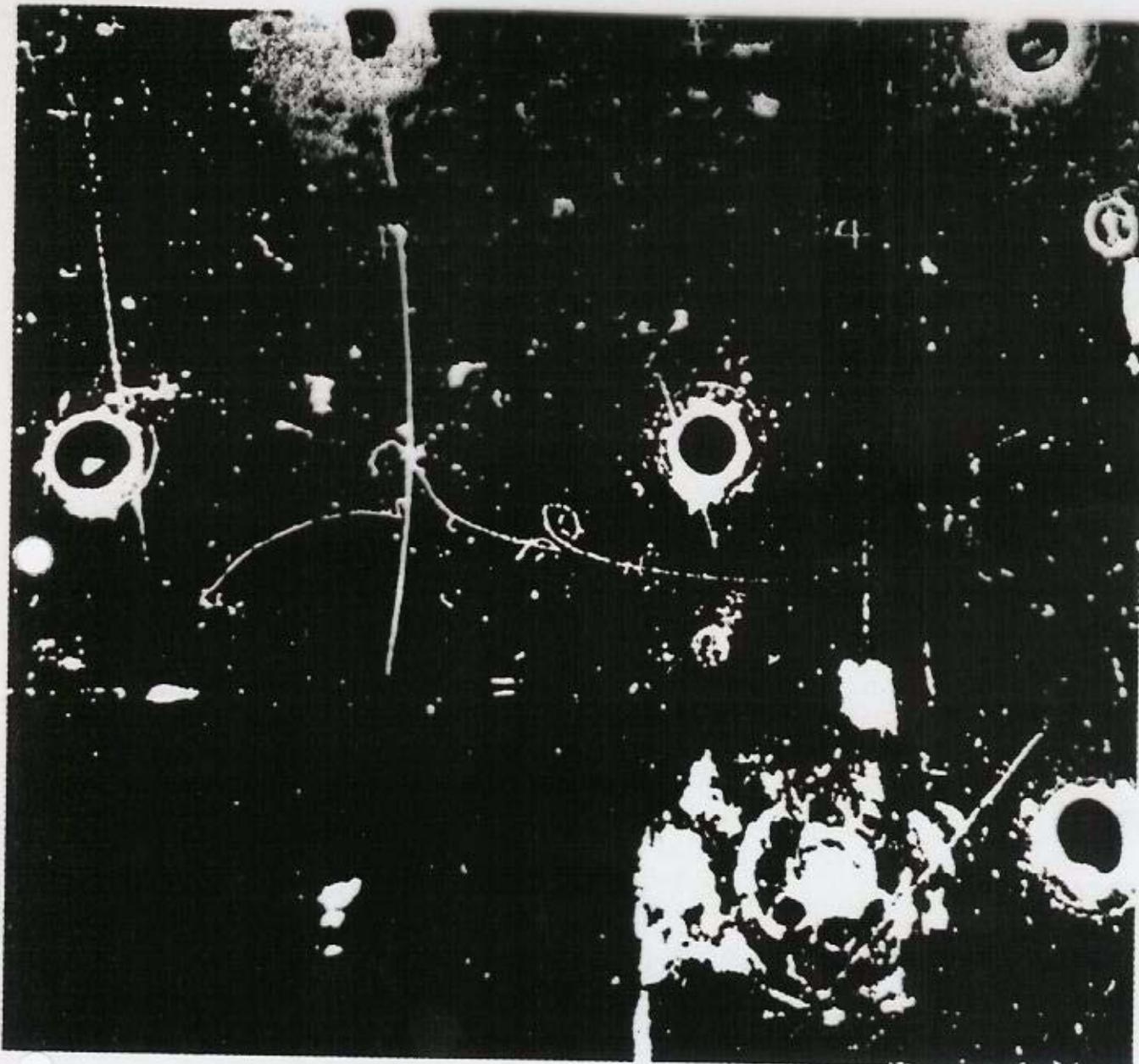
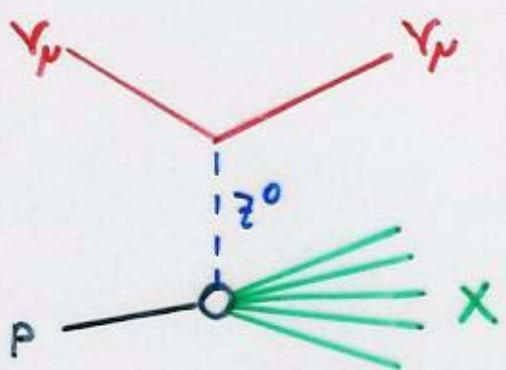
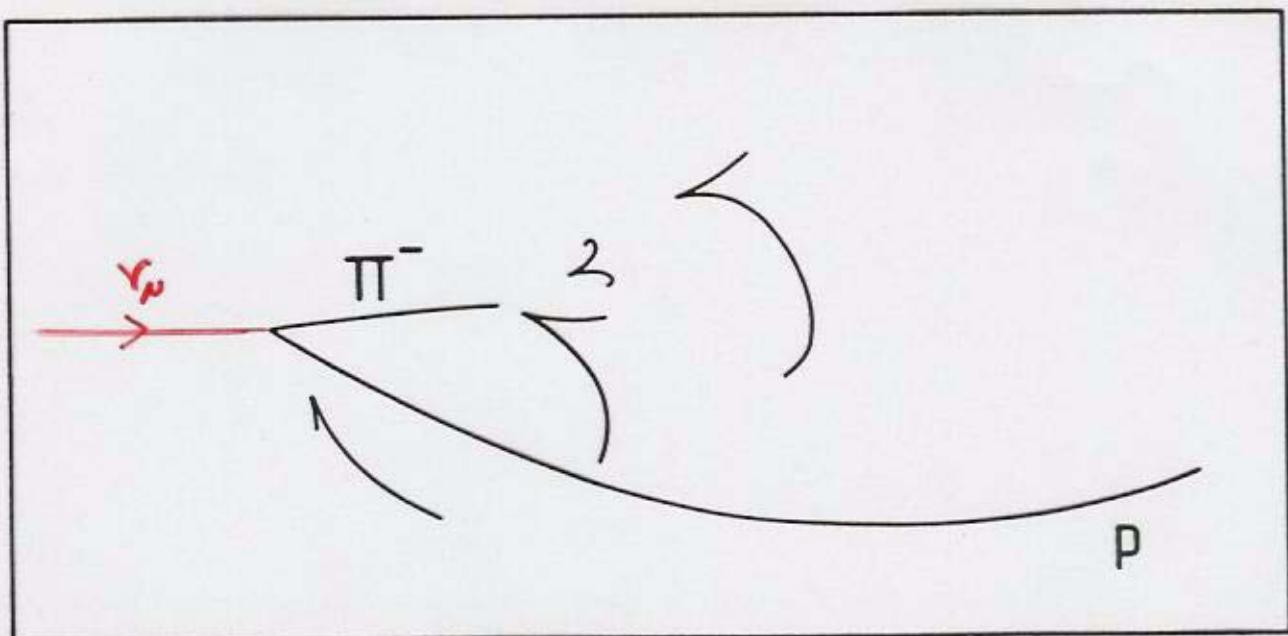
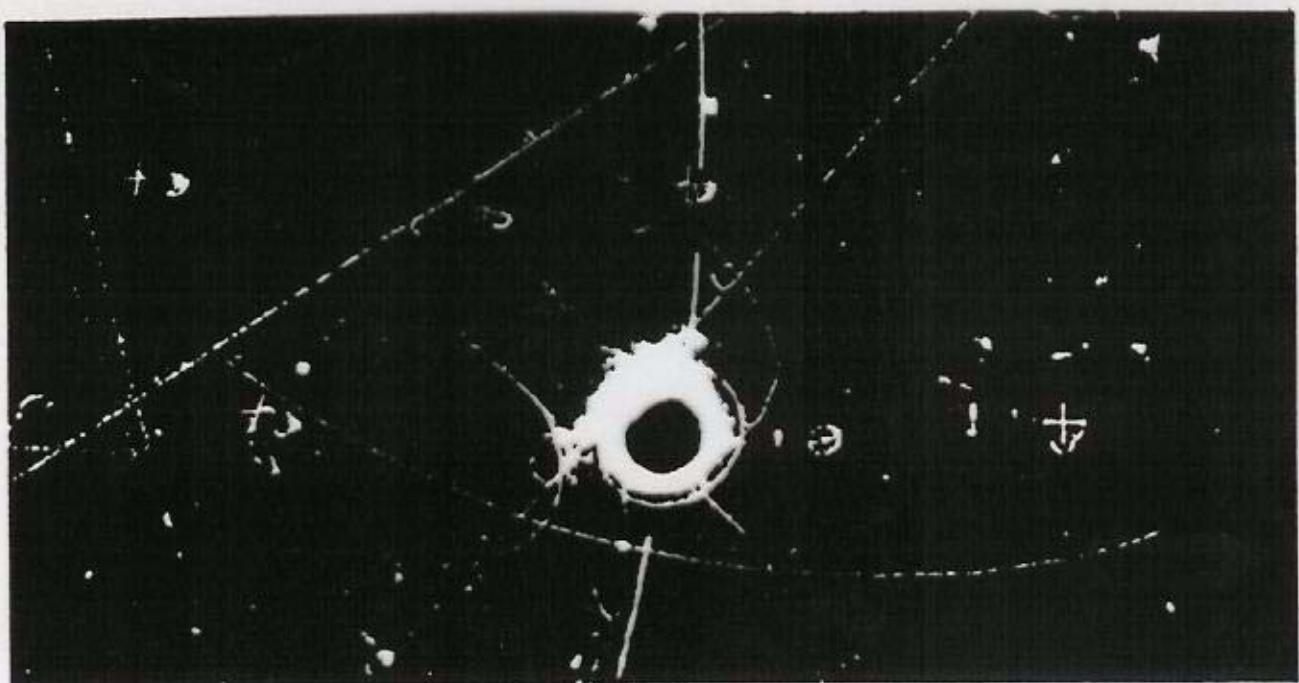


Fig. 8.4 Photograph of the first weak neutral-current event  $\bar{\nu}_\mu + e^- \rightarrow \bar{\nu}_\mu + e^- + e^+$  recorded in the Gargamelle Bubble Chamber (Häsert, 1973a). (Photograph courtesy CERN.) The annotated line diagram shows the identities of particles seen and unseen ( $\gamma$  and  $\bar{\nu}$ ).



# Uppräckten av $Z^0, W^\pm$ (1983)

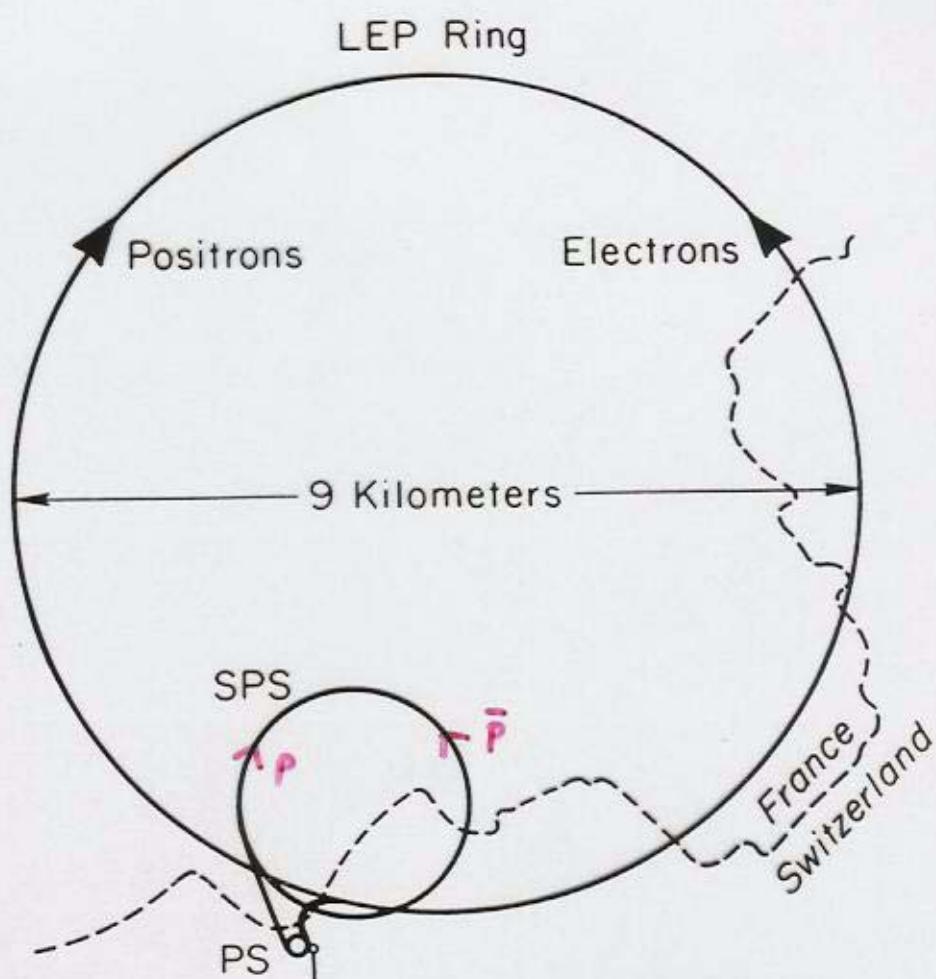
Laboratorium = CERN

Accelerator = Sp $\bar{p}$ S ( $\sqrt{s} = 540 - 900 \text{ GeV}$ )

Experiment = UA1 + UA2 (p $\bar{p}$  kollisions experiment)

Processer =  $\begin{cases} \bar{p} + p \rightarrow w^+ + x^- \rightarrow l^+ + \bar{\nu}_l + x^- \\ \bar{p} + p \rightarrow w^- + x^+ \rightarrow l^- + \bar{\nu}_l + x^+ \\ \bar{p} + p \rightarrow Z^0 + x^0 \rightarrow l^+ + l^- \end{cases}$  (där  
 $l^+ = e^+, \mu^+$   
 $l^- = e^-, \mu^-$ )

Nobelpristagare = Rubbia + Van der Meer

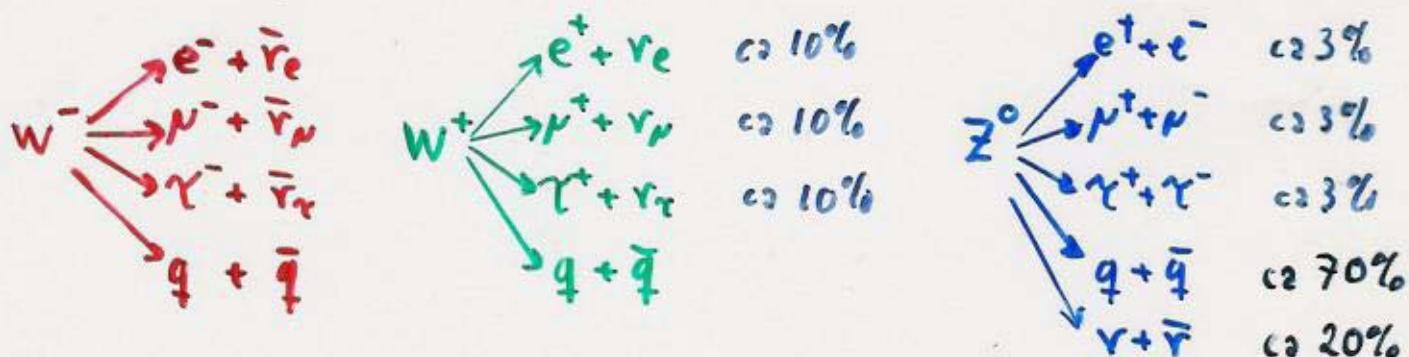


- Trärsnitt: ca 1  $Z^0$  produceras per  $10^8$   $p\bar{p}$  kollisioner  
ca 1  $W^\pm$  produceras per  $10^7$   $p\bar{p}$  kollisioner

### • Produktion:



### • Sönderfall:

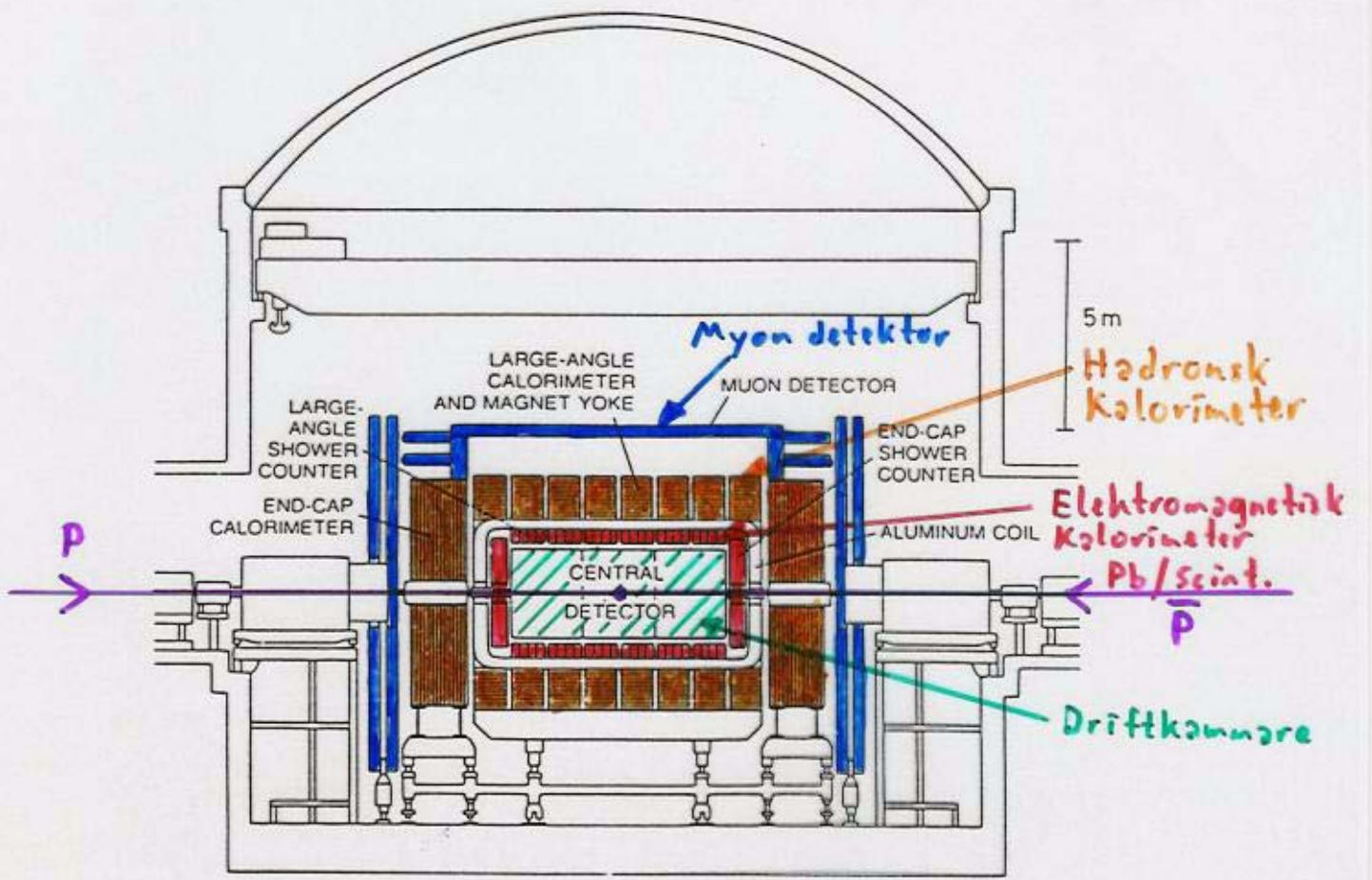
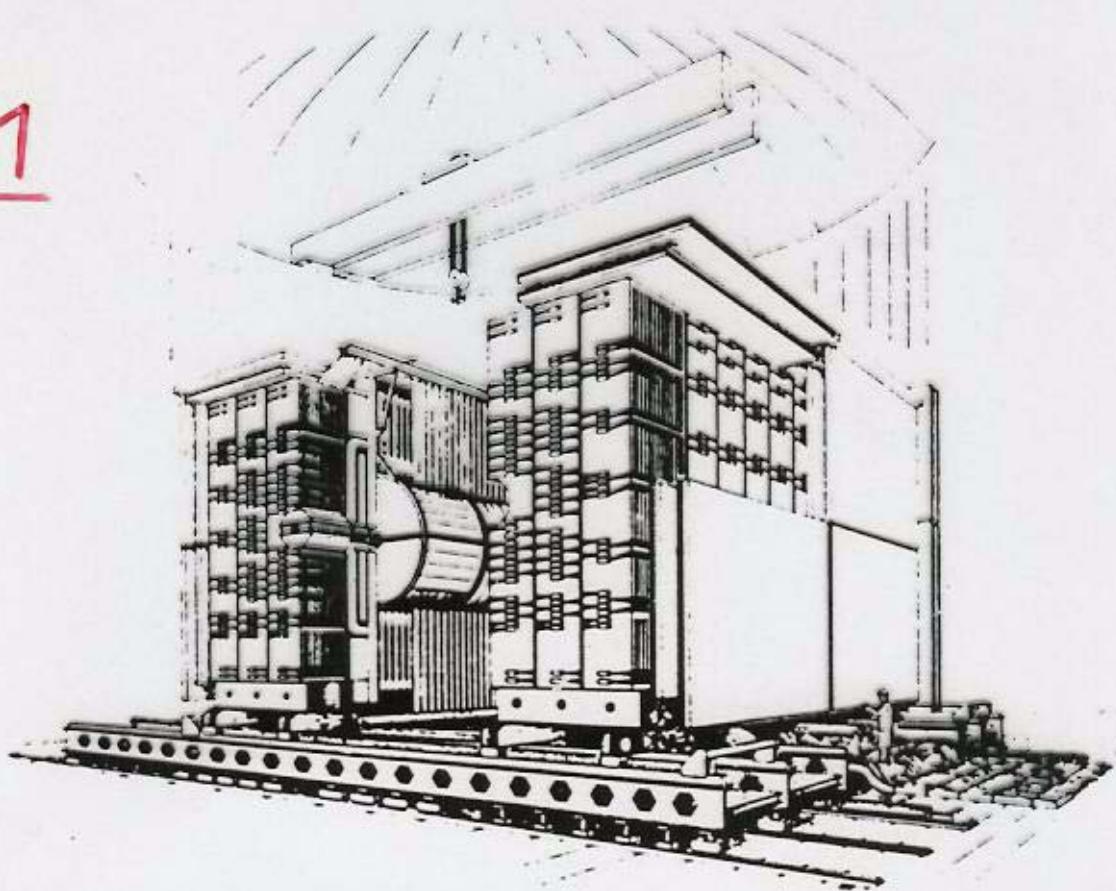


### • Massor:

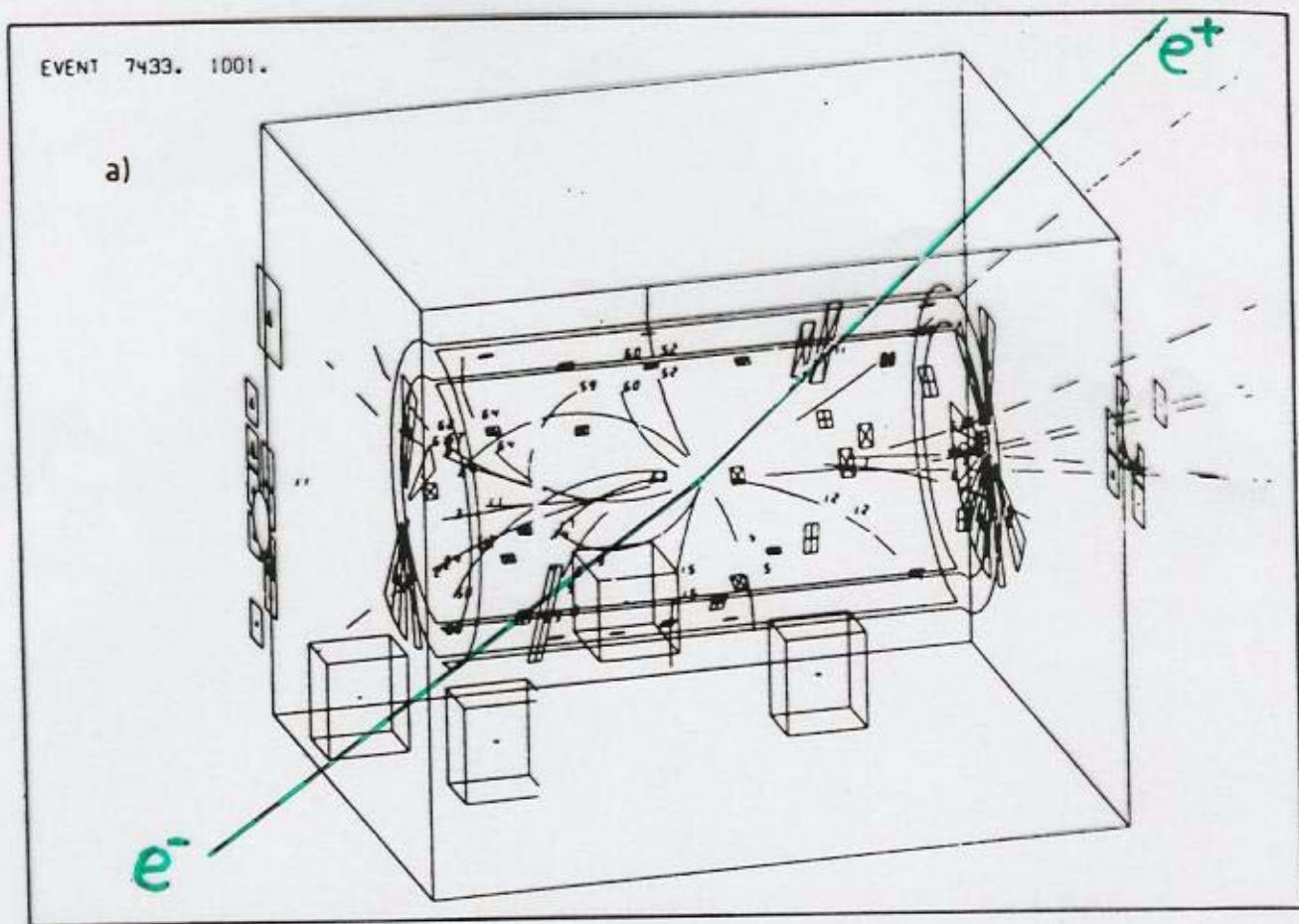
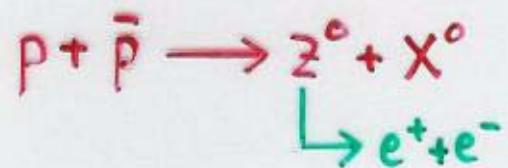
$$m_W = 81 \text{ GeV}$$

$$m_Z = 92 \text{ GeV}$$

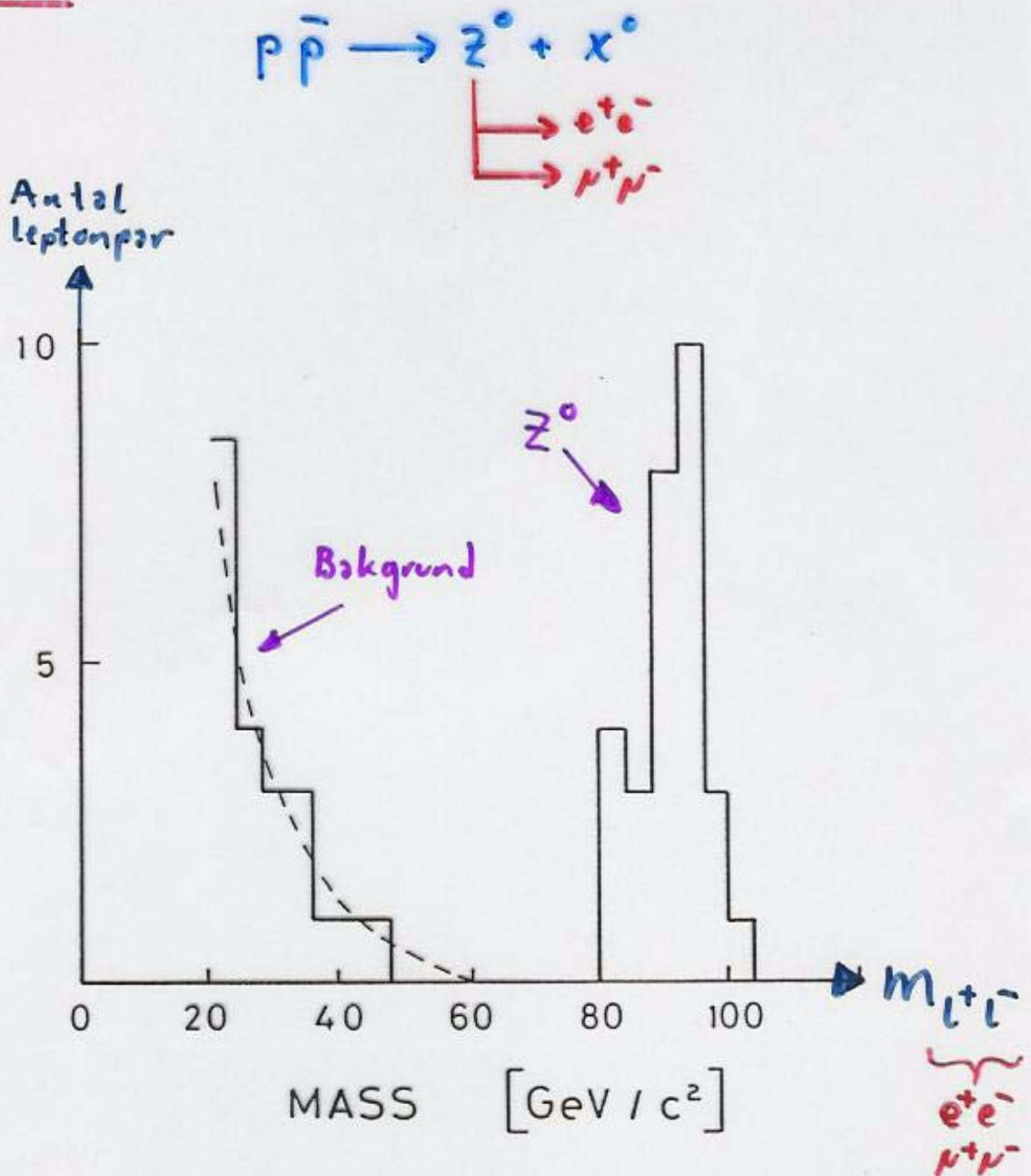
UA1



# UA1



# UA 1

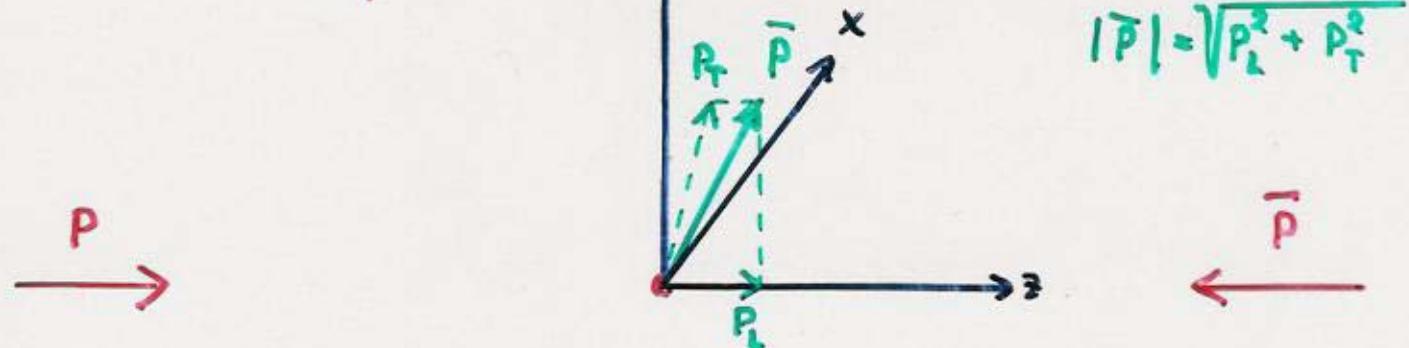


$e^- \rightarrow E_2$   
 $e^+ \rightarrow E_1$   
 $\theta$

$$m_{e^+e^-}^2 = (\vec{P}_1 + \vec{P}_2)^2 \approx 2E_1 E_2 (1 - \cos \theta)$$

W

$P_T$ :



$$|\vec{P}| = \sqrt{P_L^2 + P_T^2}$$

$P_T$ : Transversell rörelsemängd  
 $P_L$ : Longitudinell rörelsemängd

$$\left( \sum_i P_{Ti} = 0 \quad \text{där } i \text{ är partiklar} \right)$$

Massan för ett par av masslösa partiklar med öppningsvinkel  $\theta$  ges av

$$m = \sqrt{2 P_1 P_2 (1 - \cos \theta)}$$

Den transversella massan för ett par av partiklar ges av

$$m_T = \sqrt{2 P_{T1} \cdot P_{T2} (1 - \cos \theta)}$$

